



Çanakkale Onsekiz Mart University

Education Information System

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Course Information

COURSE INFORMATION

Course Title	Code	Semester	L+U Hour	Credits	ECTS
Stellar Sismology II	FZ-6026		3 + 0	3.0	7.5

Prerequisites	None
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Language of Instruction	Turkish
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Course Level	Third Cycle
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Course Type	Elective
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Mode of delivery	Face to face
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Course Coordinator	Assoc. Prof. Dr. Esin SOYDUGAN
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Instructors	Prof. Dr. Osman DEMİRKAN Prof. Dr. Ahmet ERDEM Assoc. Prof. Dr. Esin SOYDUGAN
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Assistants	
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Course Objectives	Understanding physics of pulsating stars.
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Course Content	Why do some stars pulsate? Pulsation mechanism, Period-density correlation, Pulsations in blue and red zones of instability strip. Causes for cessation of pulsation, Relation between pulsation and opacity, Obtain total internal energy of star, Equilibrium in stars (dynamic equilibrium, thermal equilibrium and vibrational equilibrium)-I, Equilibrium in stars (dynamic equilibrium, thermal equilibrium and vibrational equilibrium)-II, Lamb frequency, Brunt Väisälä frequency, Homogeneous Model and Classification of Modes, Helioseismology, Investigation of interior of Sun by Forward and Inverse methods, Determination of seismologic H-R diagram in pulsating stars like Sun, Impact of rotation and metallicity to pulsation, Physics of pulsation in Ap stars.
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Course Learning Outcomes	1) Recognize equilibrium condition in stars. 2) Recognize pulsation mechanism. 3) Have knowledge about helioseismology.
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Physics (PhD)

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Course Information

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WEEKLY COURSE CONTENT

Week	Topics	Teaching and Learning Methods and Techniques	Study Materials
1. Week	Why do some stars pulsate? Pulsation mechanism.	Oral lectures, homework, practice.	
2. Week	Period-density correlation.	Oral lectures, homework, practice.	

3. Week	Pulsations in blue and red zones of instability strip. Causes for cessation of pulsation.	Oral lectures, homework, practice.	
4. Week	Relation between pulsation and opacity.	Oral lectures, homework, practice.	
5. Week	Obtain total internal energy of star.	Oral lectures, homework, practice.	
6. Week	Equilibrium in stars (dynamic equilibrium, thermal equilibrium and vibrational equilibrium)-I.	Oral lectures, homework, practice.	
7. Week	Equilibrium in stars (dynamic equilibrium, thermal equilibrium and vibrational equilibrium)-II.	Oral lectures, homework, practice.	
8. Week	Mid-term exam.	Written exam.	
9. Week	Lamb frequency, Brunt Väisälä frequency.	Oral lectures, homework, practice.	
10. Week	Homogeneous Model and Classification of Modes.	Oral lectures, homework, practice.	
11. Week	Helioseismology.	Oral lectures, homework, practice.	
12. Week	Investigation of interior of Sun by Forward and Inverse methods.	Oral lectures, homework, practice.	
13. Week	Determination of seismologic H-R diagram in pulsating stars like Sun.	Oral lectures, homework, practice.	
14. Week	Impact of rotation and metalicity to pulsation.	Oral lectures, homework, practice.	
15. Week	Physics of pulsation in Ap stars.	Oral lectures, homework, practice.	
16. Week	Final exam.	Written exam.	

RESOURCES

Recommended Sources
-Radial and Non-radial Pulsations as Probes of Stellar Physics”, IAU Colloquium 185, Edited by Conny Aerts, Timothy R. Bedding and Jørgen Christensen-Dalsgaard, 2002, Astronomical Society of the Pacific.
-“Asteroseismology”, Conny, Aerts, 2007-2008 academic year, Lectures notes. -Copyright (2003) George W. Collins, II, http://ads.harvard.edu/books/1989fsa..book/AbookC01-C08.pdf .
-“Tidal Evolution and Oscillations in binary Stars Third Granada Workshop on Stellar Structure”, Edited by A. Claret, A. Gimenez and J.-P., Zahn, 2004, Astronomical Society of The Pacific Conference Series.

ASSESSMENT

Measurement and Evaluation Methods and Techniques		
Mid-term exam (40 percent), Final exam (60 percent).		
In-Term Studies	Quantity	Percentage
Mid Term Exam 1	1	40
Total	1	40
End-Term Studies	Quantity	Percentage

Final Exam	1	60
Total	1	60
Contribution Of In-Term Studies To Overall Grade		40
End-Term Studies		60
Total		100

COURSE CATEGORY

Course Category	Percentage
Core Courses	% 100

CONTRIBUTION OF COURSE LEARNING OUTCOMES TO PROGRAMME OUTCOMES

Programme Outcomes	Contribution Level	DK1	DK2	DK3
<u>PY1</u>	5	4	5	5
<u>PY2</u>	4	4	4	3
<u>PY3</u>	4	3	4	4
<u>PY4</u>	4	4	4	3
<u>PY5</u>	4	4	3	4
<u>PY6</u>	5	5	5	4
<u>PY7</u>	4	4	4	3
<u>PY8</u>	4	4	3	4
<u>PY9</u>	5	5	5	4
<u>PY10</u>	3	3	3	3
<u>PY11</u>	4	4	4	3
<u>PY12</u>	4	4	4	3
<u>PY13</u>	4	4	4	3
<u>PY14</u>	5	4	5	5
<u>PY15</u>	5	5	5	4

*DK = Course's Contribution.

	0	1	2	3	4	5
Level of contribution	None	Very Low	Low	Fair	High	Very High

ECTS CREDITS AND COURSE WORKLOAD

Event	Quantity	Duration (Hour)	Total Workload (Hour)
Class Hours (14 weeks)	14	3	42
Final Exam Preparation	1	36	36
Mid Term Exam Preparation	1	32	32
Further Study	14	1	14
Final Exam	1	3	3
Mid Term Exam 1	1	3	3
Laboratory	2	10	20
Preliminary Study	14	3	42

Total Workload	192
Total Workload / 25.5 (s)	7.53
ECTS Credit of the Course	8

